

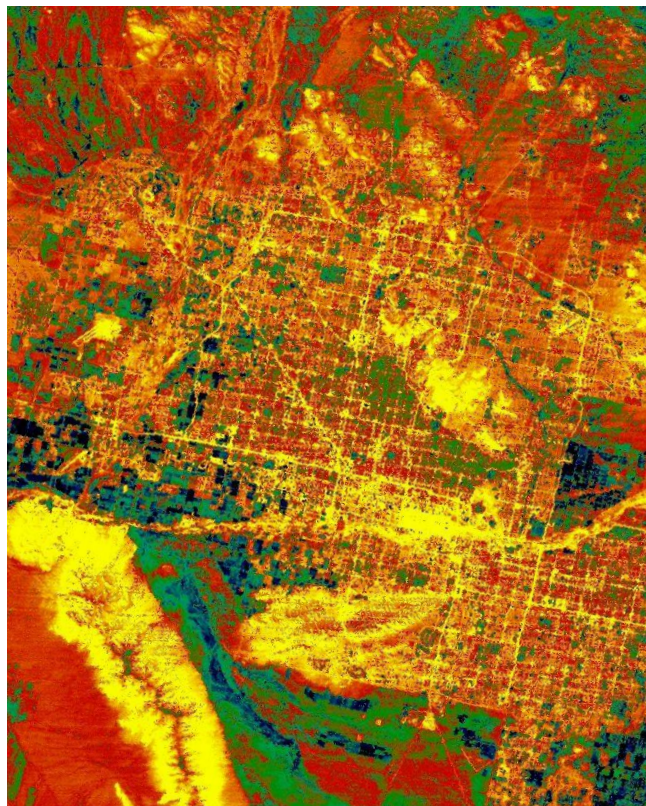
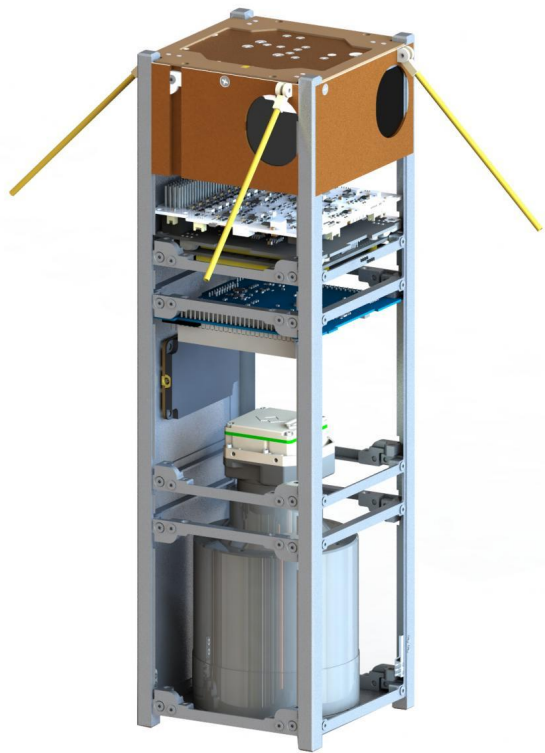


# Phoenix CubeSat: Radiometric Work on the Payload Camera

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Presented by Andre De Simone and Daniel La Rosa

# Overview

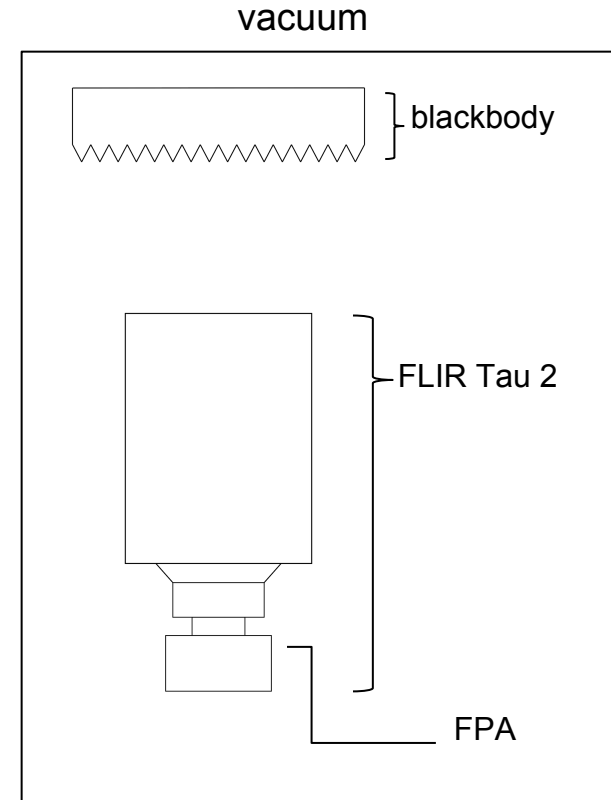


# Thermal Vacuum Test

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# Thermal Vacuum Test - Overview

- A focal-plane array (FPA) is a sensor that detects infrared wavelengths and converts them into images.
- Differences in FPA temperature may affect image value accuracy.
- It is crucial to find the camera's optimal operating temperature so that it will produce meaningful data.
- Optimal FPA temperature can be found by pointing the camera at a blackbody and varying temperature while inside a vacuum chamber.



# Thermal Vacuum Test - Setup

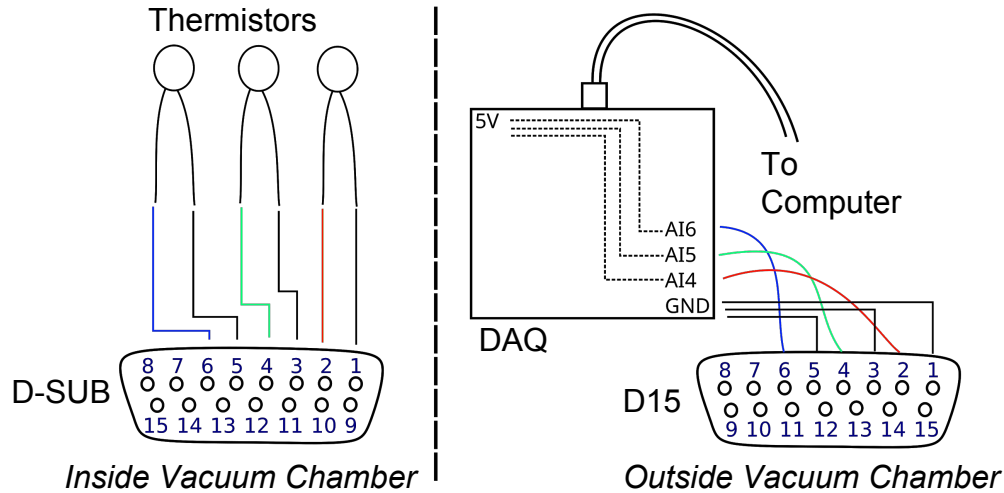


Figure 1. Diagram of the thermistor setup.

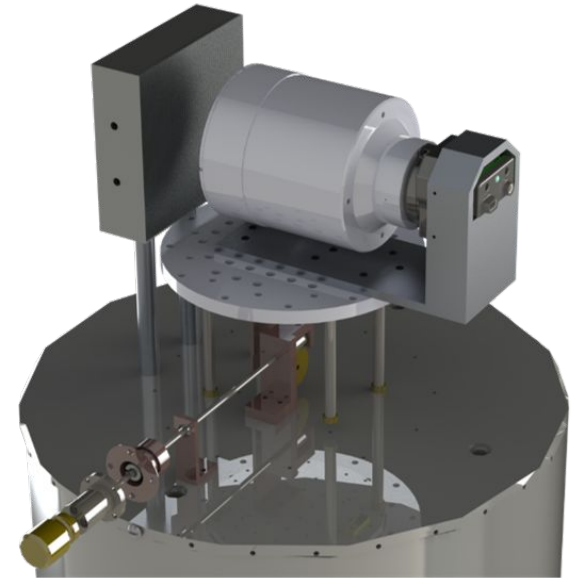
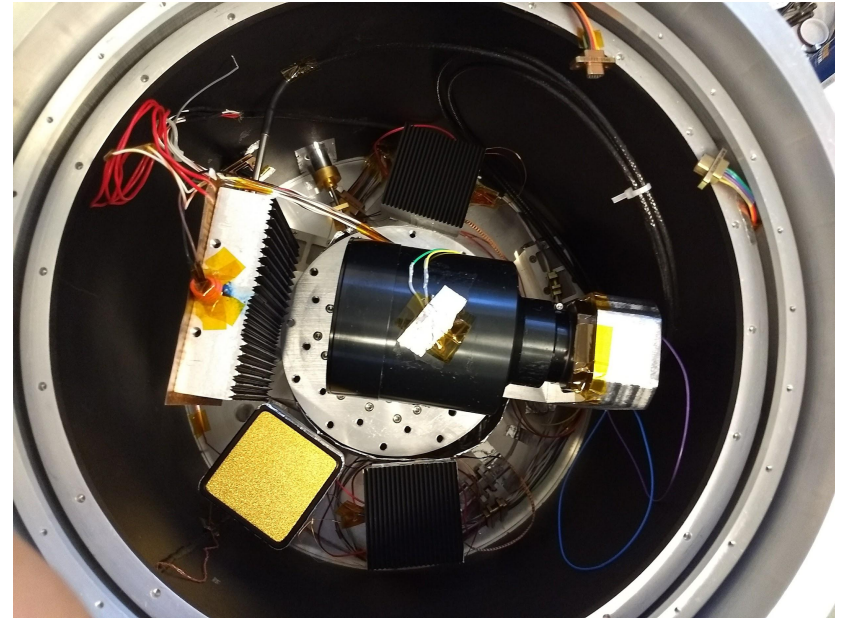
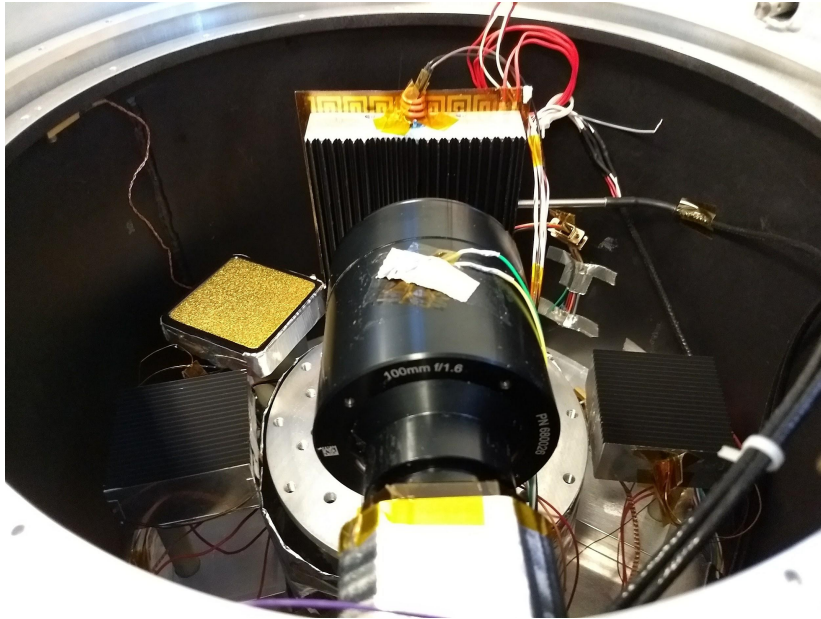
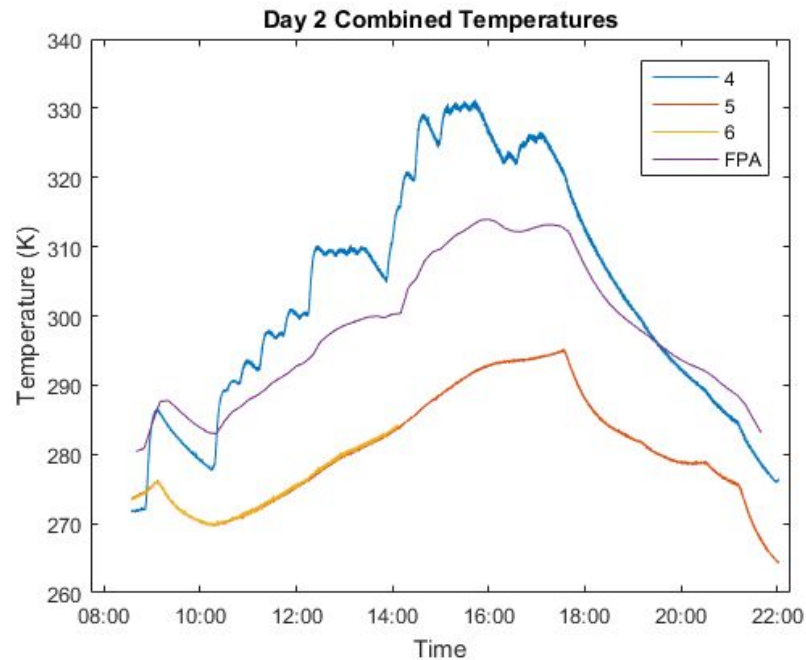
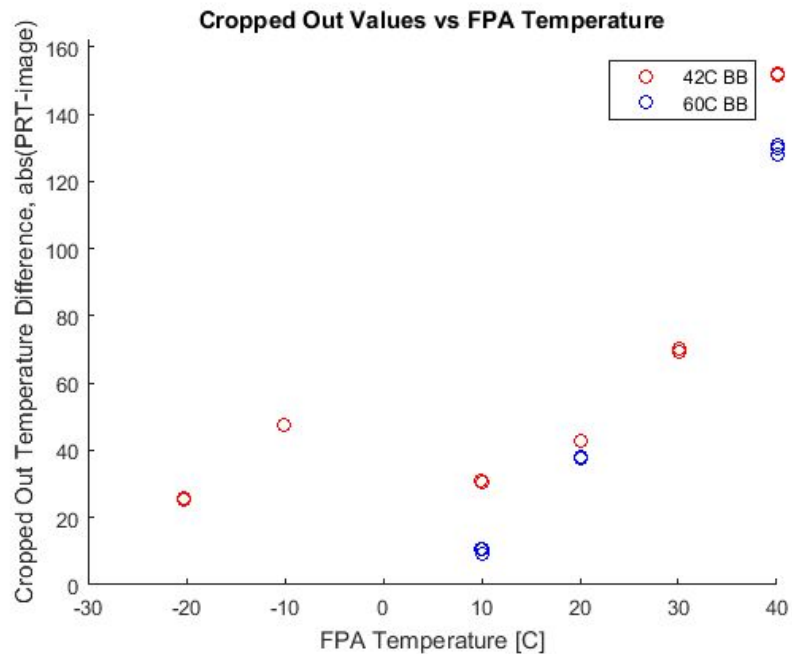


Figure 2. T-VAC test setup CAD model.

# Thermal Vacuum Test - Setup (cont.)



# Thermal Vacuum Test - Results

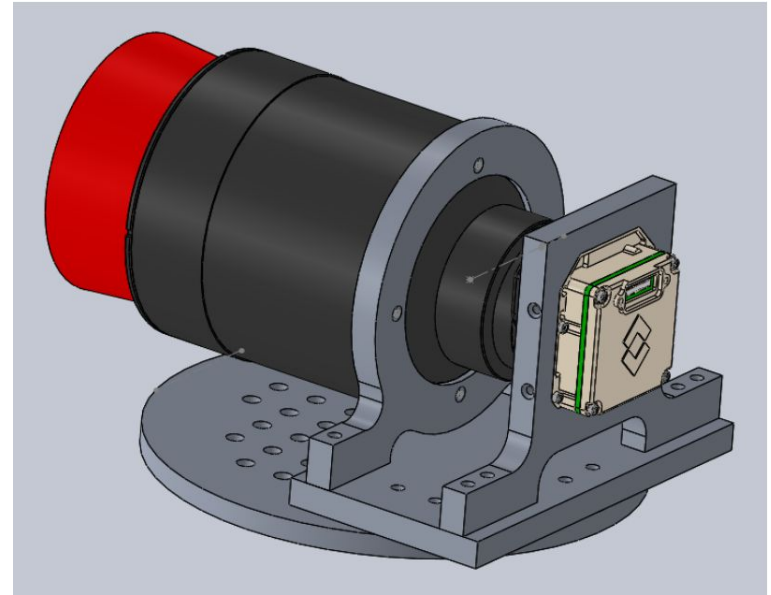
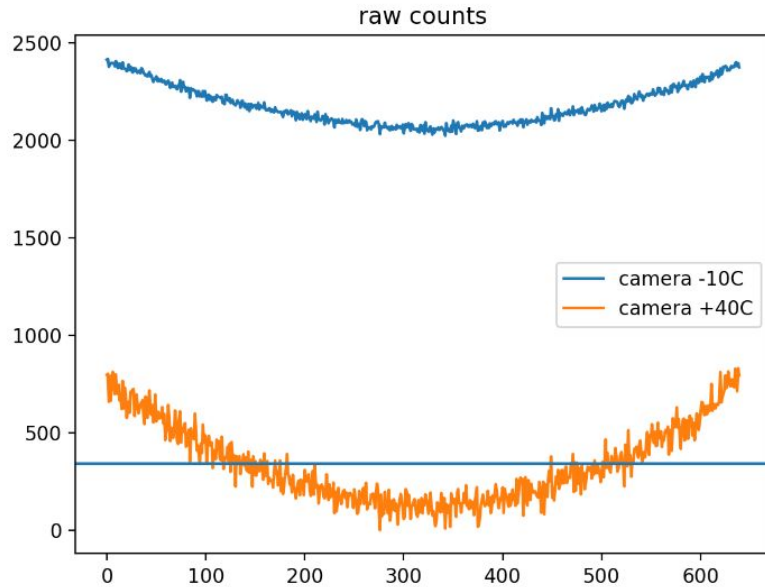


# Future Work

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# Thermal Vacuum Test - Possible Avenues of Error



# Lunar Imaging

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# Lunar Imaging Test - Overview

- Goal: To image the moon and come up with an atmospheric model to accurately measure surface temperature.
- Reasoning: To try to recreate the conditions the camera will experience during flight.
  - Image the moon at its zenith over Tempe, AZ



# Lunar Imaging Test - Setup

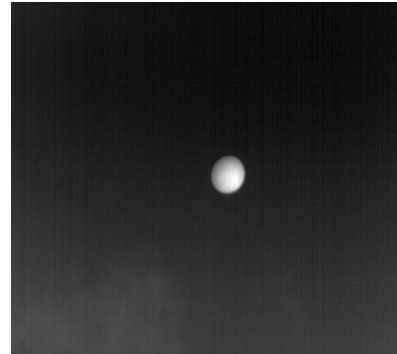
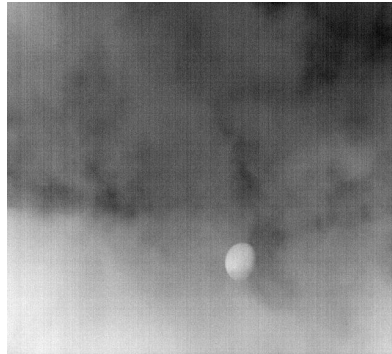
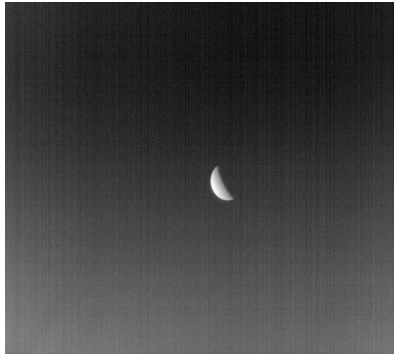
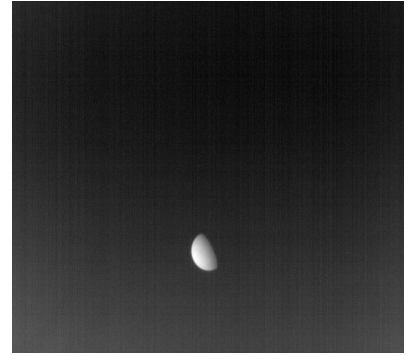
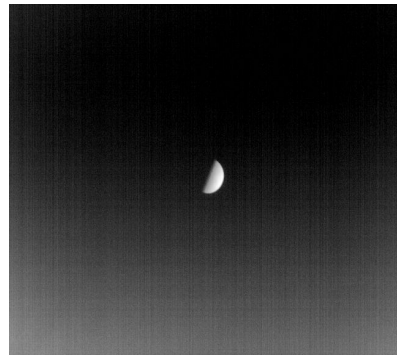
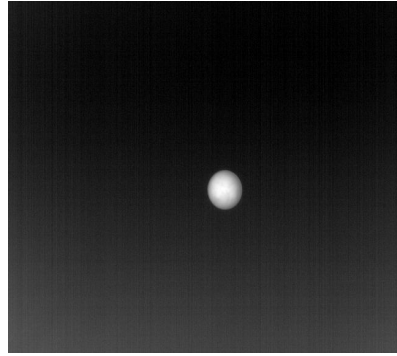
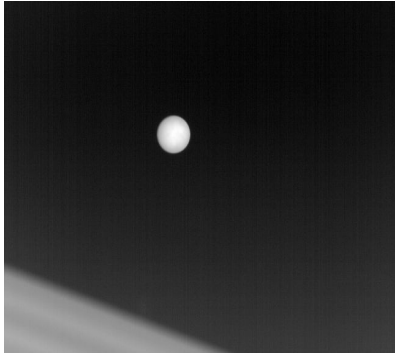
- Tripod mount with 3 axis alligator clamp



# Lunar Imaging Test - Setup

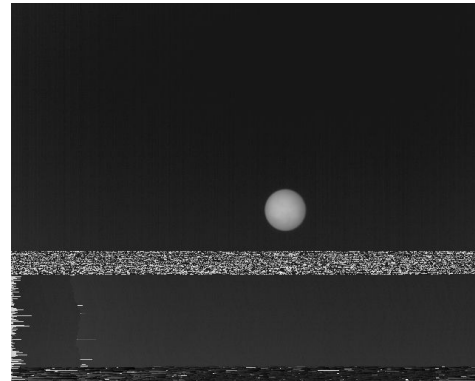
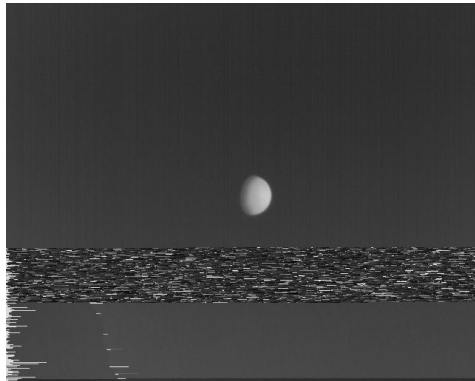
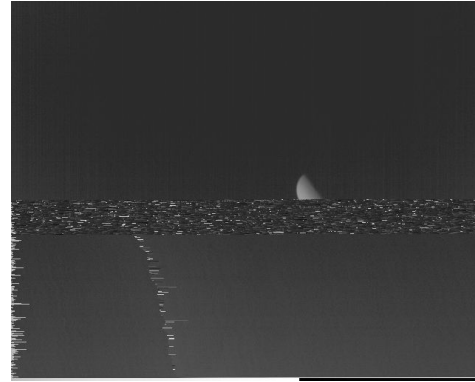
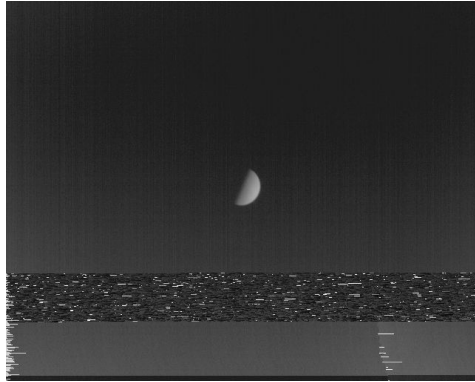


# Lunar Imaging Test - Results

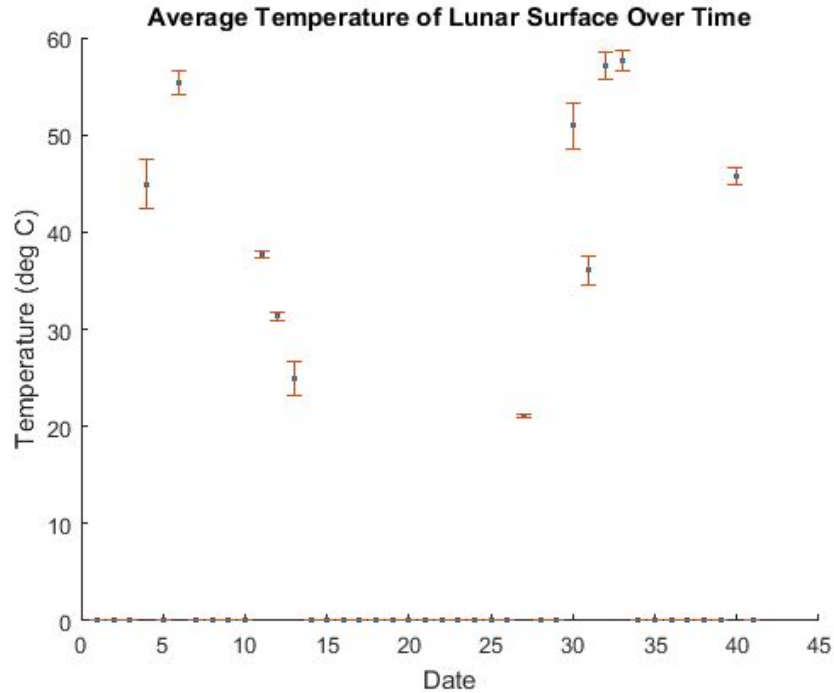




# Lunar Imaging Test - Results



# Lunar Imaging Test - Results



Start date corresponds Oct 1st, end day corresponds to Nov 10th



# Future Work

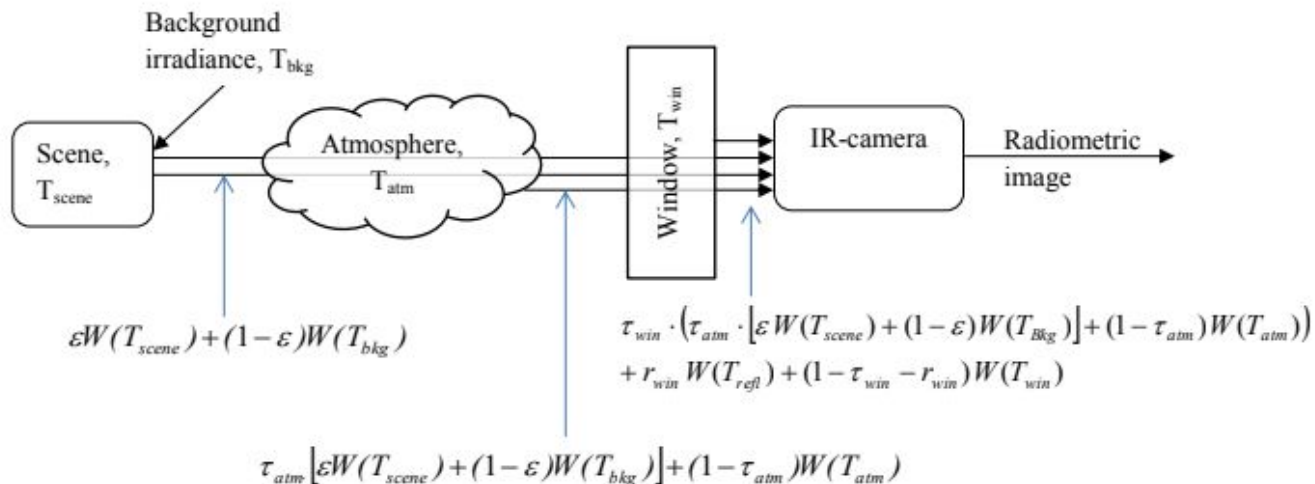
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# Lunar Imaging/Atmospheric Modeling

- After addressing issues with thermal chamber testing, lunar imaging will be revisited for accurate modeling of atmosphere
- More data will be taken
- Will be tested with filter during cloudy day
  - 10.5-12.5 $\mu\text{m}$



# Manipulating the Raw Data from the Camera



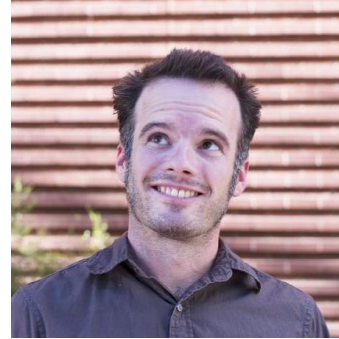
The incident radiation onto the camera is given by

$$S = \tau_{win} \cdot (\tau_{atm} \cdot [\epsilon W(T_{scene}) + (1 - \epsilon)W(T_{bkg})] + (1 - \tau_{atm})W(T_{atm})) + r_{win} W(T_{refl}) + (1 - \tau_{win} - r_{win})W(T_{win})$$

# Acknowledgements



- **Judd Bowman**
- *Mentor*
- *School of Earth and Space Exploration*



- **Danny Jacobs**
- *Mentor*
- *School of Earth and Space Exploration*



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Questions?

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